Thanks to George & Baha for Their Interest & Support in Sterile Neutrinos Over the Past 30 Years!

- I first met George at a Gordon Conference in 1995
- And I first met Baha shortly later at ORNL
- I blame them both for whetting my interest in sterile neutrinos!





Evidence for Sterile Neutrinos

W. C. Louis, January 17, 2025

- Sterile Neutrino Evidence (> 3σ)
- Sterile Neutrino Indications/Hints (~2 σ) & Tensions in the Data
- Quasi-sterile Neutrinos & Apparent CPT Violation
- Upcoming Experiments
- Conclusions

Sterile Neutrino Evidence (> 3σ)

- LSND
- MiniBooNE
- Gallium Anomaly



LSND Liquid Cherenkov/Scintillation Detector (1993-1998)

Located 30m from the beam dump

167 tons of Mineral Oil (CH_2) with 0.031 g/l of b-PBD & 1240 8" PMTs

Events are reconstructed from the charge/time of each hit PMT

Designed to search for $\overline{\nu}_{\mu} \rightarrow \overline{\nu}_{e}$

 $\overline{v}_e p \rightarrow e^+ n$, $n p \rightarrow d \gamma (2.2 \text{ MeV})$

LSND Event Excess

A. Aguilar et al., Phys. Rev. D 64, 112007, (2001)

Correlated γ = 117.9+-22.4 events Excess = 87.9+-22.4+-6.0 events



 $\overline{v}_e p \rightarrow e^+ n$, $n p \rightarrow d \gamma (2.2 \text{ MeV})$

Selection	On Even	ts Off Events	v Backgrou	nd Excess	Significance
R _γ > 1	205	106.8+-2.5	39.2+-3.1	59.0+-12.7	4.6σ
R _γ > 10	86	36.9+-2.3	16.9+-2.3	32.2+-8.0	4.0σ
R _γ > 100	27	8.3+-0.7	5.4+-1.0	13.3+-3.9	3.4σ

LSND collected 1.9x10²³ (0.3g) protons on target and observed a 3.8 σ excess of events. Consistent with $\overline{v}_{\mu} \rightarrow \overline{v}_{e}$ oscillations, corresponding to P_{osc} = (0.264+-0.067+-0.045)%.

LSND Event Excess

A. Aguilar et al., Phys. Rev. D 64, 112007, (2001)



The LSND best fit at 1.2 eV² is favored over null by ~5 σ ($\Delta \chi^2 = 29$). LSND had very low intrinsic $\overline{\nu_e}$ background (0.07%) and is hardly affected by $\overline{\nu_e}$ disappearance.

Joint LSND/KARMEN Analysis

E. D. Church, K. Eitel, G. B. Mills, and M. Steidl, Phys. Rev. D66, 013001, (2002)



MiniBooNE Experiment



- Similar L/E as LSND for $v_{\mu} \rightarrow v_e \& \overline{v_{\mu}} \rightarrow \overline{v_e}$ oscillations
 - MiniBooNE ~500m/~500MeV
 - LSND ~30m/~30MeV
- Horn focused neutrino beam (p+Be)
 - Horn polarity → neutrino or anti-neutrino mode
- 800t mineral oil Cherenkov detector

Total Neutrino Data 18.75 x 10²⁰ POT

Phys.Rev.D 103, 052002 (2021)



Neutrino Excess = 560.6 + 119.6 events (4.7 σ) with 200<E<1250 MeV

Excess is larger than LSND at low energies & high $\cos\theta$

Total Neutrino Data

Phys.Rev.D 103, 052002 (2021)



Neutrino Excess = 560.6 +- 119.6 events (4.7σ) with 200<E<1250 MeV

Allowed Regions



Phys. Rev.D 103, 052002 (2021)

Neutrino + Anti-Neutrino Mode

 $(\Delta m^2, \sin^2 2\theta) = (0.043 \text{ eV}^2, 0.807)$ $\chi^2/ndf = 21.7/15.5 \text{ (prob = 12.3\%)}$

Total Excess = 638.0 +- 132.8 events (4.8σ) with 200<E<1250 MeV



A 3.4 Mci 51Cr source was placed inside a vat containing 40 tons of Ga - 69Ga (0.6) & 71Ga (0.4) The 51Cr source ($\tau_{1/2}$ = 11.43 d) produces 0.7 MeV v_e that interact with 71Ga to produce 71Ge The 71Ge atoms are extracted periodically and counted







The Gallium experiments measure a relative event rate = 0.80+-0.05 (4 σ deficit!) The best fit for the Gallium anomaly is $\sin^2 2\theta = 0.34$, $\Delta m^2 = 1.25 \text{ eV}^2$



FIG. 10. Allowed regions for two BEST results. The best-fit point is $\sin^2 2\theta = 0.42^{+0.15}_{-0.17}$, $\Delta m^2 = 3.3^{+\infty}_{-2.3} \text{ eV}^2$ and is indicated by a point.

FIG. 12. Allowed regions for two GALLEX, two SAGE and two BEST results. The best-fit point is $\sin^2 2\theta = 0.34^{+0.14}_{-0.09}$, $\Delta m^2 = 1.25^{+\infty}_{-0.25}$ eV² and is indicated by a point.

The Gallium experiments have the **best** determination of the absolute event rate of any neutrino experiment!

Comparison of LSND, MB, & Gallium Anomaly

- LSND 3.8 σ Sensitive almost entirely to $v_{\mu} \rightarrow v_{e}$
- MiniBooNE 4.8 σ Sensitive to both $v_{\mu} \rightarrow v_{e}$ and $v_{e} \rightarrow v_{x}$
- Gallium 4.0 σ Sensitive entirely to $v_e \rightarrow v_x$

The signals from these experiments are all model independent! (They do not depend, e.g., on a 3+1 model)

Sterile Neutrino Indications/Hints (~ 2σ) & Tensions

- IceCube vs MINOS+
- MicroBooNE vs MiniBooNE
- Reactors vs Gallium Anomaly

Indication of v_{μ} Disappearance from IceCube Phys. Rev. Lett. 133, 201804 (2024)

IceCube is sensitive to \overline{v}_{μ} disappearance from matter effects at an $L/E \sim 1!$

IceCube does observe a closed 95% contour consistent with disappearance.

This is the first hint of v_{μ} disappearance at short baseline.

Upcoming results from IceCube will have lower systematic uncertainties and better sensitivity.

Best fit at Δm^2 = 3.5 eV², sin²2 θ = 0.16



Indication of v_{μ} Disappearance from IceCube Phys. Rev. Lett. 133, 201804 (2024)

However, IceCube (\overline{v}) is in strong tension with MINOS+ (v)

The MINOS+ limit is 2x better than the sensitivity; Is the limit too good?

Apparent CPT Violation?



MicroBooNE Observes a ν_e Deficit Consistent with the Gallium Anomaly!



WireCell Reconstruction

The Elephant in the room!



MicroBooNE Observes a v_e Deficit Consistent with the Gallium Anomaly!





- Is some of the MiniBooNE excess due to BSM photon or e⁺e⁻ production? (but antineutrino excess should be higher)
- Is difference due to different distances for MicroBooNE (470m) & MiniBooNE (541m)?
- Data suggest a more complicated 3+N model with uB and MB observing both v_e appearance and v_e disappearance with uB dominated by v_e disappearance and MB dominated by v_e appearance.

Reactor Neutrino Anomaly Phys.Lett.B 829 (2022)



- Reactor v Experiments measure fewer events than expected.
- However, the systematic uncertainties are large.
- The deficits are not significant.

Difference Between Reactor & Gallium Anomalies

arXiv:2209.00916 3+N Model? Apparent CPT Violation?



Neutrino 4 Evidence for Sterile Neutrinos

arXiv:2005.05301

- Neutrino-4 observes wiggles, Corresponding to $\Delta m^2 = 7.3 \text{ eV}^2 \& \sin^2 2\theta = 0.39 \& \text{ consistent with the}$ Gallium anomaly
- Are these wiggles due to statistics or to ν oscillations?
- However, taking into account energy resolution, $\sin^2 2\theta \sim 1$



Quasi-sterile Neutrinos & Apparent CPT Violation

• BSM Matter Effects can cause apparent CPT violation

Quasi-sterile neutrinos from dark sectors. Part I. BSM matter effects in neutrino oscillations and the short-baseline anomalies.

Daniele S. M. Alves, WCL, and Patrick G. deNiverville JEHP 08, 034 (2022)



Figure 1. A generic consequence of dark sectors with neutrino- and vector-portals are quasi-sterile neutrinos. The BSM matter potential generated by the dark vector interactions can alter neutrino oscillations in matter.

Quasi-sterile neutrinos from dark sectors. Part I. BSM matter effects in neutrino oscillations and the short-baseline anomalies.

Daniele S. M. Alves, William C. Louis and Patrick G. deNiverville JEHP 08, 034 (2022)

Maximal neutrino oscillations occur at the resonant neutrino energy!

However, there is no resonance for antineutrinos.



Quasi-sterile neutrinos from dark sectors. Part I. BSM matter effects in neutrino oscillations and the short-baseline anomalies.

Daniele S. M. Alves, William C. Louis and Patrick G. deNiverville JEHP 08, 034 (2022)

The Quasi-sterile v model explains the MiniBooNE excess well & is consistent with MicroBoone.

The Quasi-sterile v model predicts a strong distance dependence near the resonance.

This fit is not optimized!



Upcoming Experiments

- SBN
- JSNS²
- KATRIN



SHORT-BASELINE NEUTRINO PROGRAM

Short Baseline Neutrino (SBN) Program



MicroBooNE took data from 2015-2020 ICARUS started data taking in 2021 SBND began data taking in 2024

Short Baseline Neutrino (SBN) Program Has Two v Beams



SBN has 3 LAr TPC detectors and 2 beamline (BNB & NuMI) This yields 5 different baselines: SBND: 110m (BNB) ICARUS: 600m (BNB) & 810m (NuMI) MicroBooNE: 470m (BNB) & 680m (NuMI)

ICARUS

Four TPCs in two modules Each module is 3m x 4m x 18m Steady data taking since March 2021



ICARUS has Cosmic Ray Tracker (CRT) & concrete overburden



SBND



Membrane cryostat Completed in October 2022



Detector installed in December 2022 Two Time Projection Chambers with Total dimension: 4m x 4m x 5m

APA – Anode wire planes with ~11000 wires CPA - Cathode covered with TPB coated reflectors Photon Detection System: 120 PMTs, 192 X-Arapucas

SBN Will have 5σ Coverage of the Short-Baseline v Anomalies!







JSNS2 will provide a direct test of the LSND Signal J-PARC MLF facility: 25 Hz, 1 MW, 3 GeV RCS Each pulse consists of 2x100 ns buckets, 600 ns apart



50t (Gd-loaded + unloaded) liquid scintillator detector (4.6m diameter x 4.0m height) 120 10" PMTs

JSNS² vs. LSND

	LSND	JSNS ²	Notes
Detector Mass	167t	17t	-
Baseline	30m	24m	-
Beam Proton Energy	0.8GeV	3GeV	Allows for KDAR measurement. Expect ~10x higher pion production
Beam Power	800kW	1MW	-
Beam Duty Factor	600µs x120Hz	100ns(x2) x25Hz	Expect ~300x fewer ambient IBD backgrounds
Detector Medium	Dilute LS	Gd-LS	-
Neutron Capture	Hydrogen, ~0.2ms, 2.2MeV	Gadolinium, ~26µs, 8MeV	Shorter capture time & higher energy mean fewer backgrounds
Particle ID	Cherenkov	9 PSD	-

JSNS² Will Cover the LSND Signal at 3σ



Near Detector: 17 tons, 120 10-inch PMTs, 24 m Far Detector: 32 tons, 220 10-inch PMTs, 48 m

KATRIN Tritium Beta Decay Experiment



KATRIN has also set limits on the mass of v_4 which is mostly sterile.

KATRIN has sensitivity to fully cover the Neutrino 4 signal.

Nexus of $\beta\beta$ decay, tritium β decay, reactors, & radioactive source expts.

KATRIN has set a limit on the v_1 mass of ~0.45 eV, with an ultimate sensitivity of <0.3 eV.



KATRIN Tritium Beta Decay Experiment



KATRIN has set a limit on the v_1 mass of ~0.45 eV, with an ultimate sensitivity of <0.3 eV.

KATRIN can also set limits on the mass of v_4 which is mostly sterile.

KATRIN has sensitivity to fully cover the Neutrino 4 signal



Conclusions

- Strong evidence for sterile neutrinos, although the 3+1 model is probably too simplistic
- Sterile neutrinos are great dark matter candidates and could help explain the Hubble Constant discrepancy
- Upcoming experiments (SBN, JSNS², KATRIN) may be able to prove overwhelmingly the existence of sterile neutrinos

MicroBooNE Observes a ν_e Deficit Consistent with the Gallium Anomaly!

Number of events

	Signal cl	Signal channel	
	$1eNp0\pi$	$1e0p0\pi$	
Data counts	102	41	
Total H_0 prediction (constrained)	133.5 ± 7.4	47.6 ± 3.7	
Total H_1 prediction (constrained, signal model 1)	151.7 ± 8.2	55.3 ± 3.9	
Total H_1 prediction (constrained, signal model 2)	168.5 ± 9.3	57.7 ± 4.1	
Total H_0 prediction (unconstrained)	123.6 ± 17.1	42.8 ± 7.6	
Cosmics (unconstrained)	0.8	5.3	
$\nu_e \ \mathrm{CC} \ (\mathrm{unconstrained})$	104.1	18.4	
ν other (unconstrained)	6.8	6.1	
$\nu \text{ NC } \pi^0 \text{ (unconstrained)}$	11.9	13.1	
MiniBooNE LEE Signal Model 1 (unconstrained)	13.4	5.0	
MiniBooNE LEE Signal Model 2 (unconstrained)	29.9	7.5	



G. Cerati (uB, FNAL) - Fermilab "Wine&Cheese" seminar

ICARUS



SBND





Membrane cryostat constructed inside an outer warm steel structure

Central cathode with two 2m drift regions

Two 4m x 2.5m Anode Plane Assemblies (APAs) per side (drift region)